

# Non-Equilibrium Ionization State and Two-Temperature Structure in the Linked Region of Abell 399/401

...will be submitted to PASJ



@Tsukuba

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The Warm & Hot Universe

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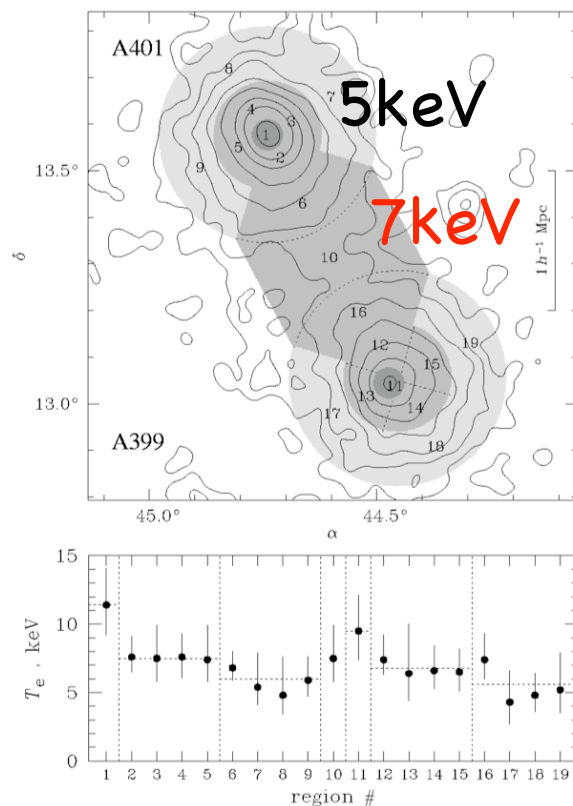
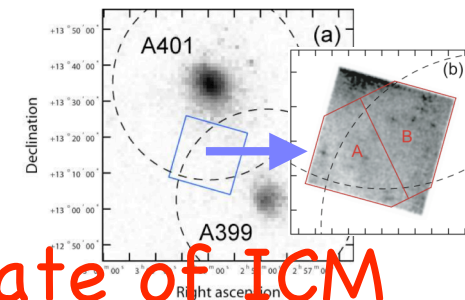


# 1) Introduction to Abell 399/401

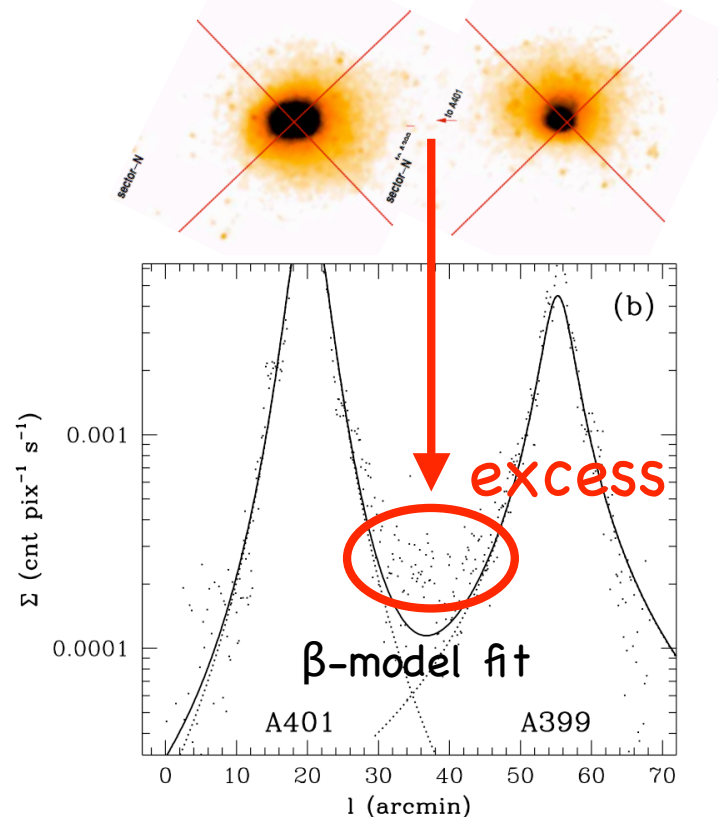


- On an early stage of a merging
- High Metallicity in the linked region

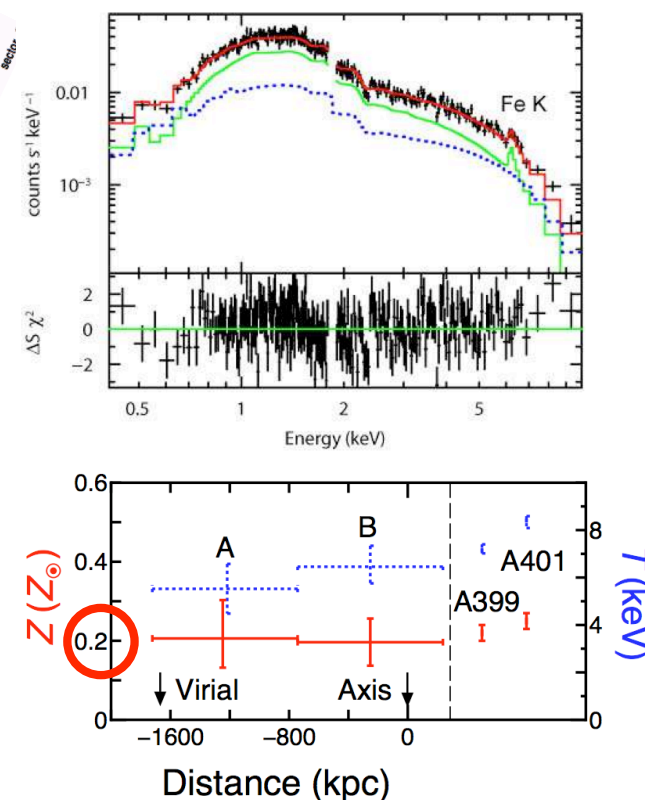
↑ assuming the "Equilibrium" state of ICM



ASCA: Temp. map  
Markevitch+ '98



XMM-Newton: Bright. profile  
Sakelliou, Ponman '04



Suzaku: Metallicity  
Fujita+ '08



## 2) Equilibrium Timescale



- The linked region → Gas density  $\sim 10^{-4} \text{ (cm}^{-3}\text{)}$ 
  - Timescale of Collisional ionization equilibrium
    - $t_{\text{ion-rec}} \sim 3 \times 10^{-2} \text{ Gyr } (n/10^{-2} \text{ cm}^{-3})^{-1}$  (e.g., Masai '84)
  - Timescale of e-i temperature equilibration
    - $t_{\text{e-i}} \sim 2 \times 10^{-2} \text{ Gyr } (n/10^{-2} \text{ cm}^{-3})^{-1} (T/10^8 \text{ K})$  (Spitzer '56)

Area · Density ( $\text{cm}^{-3}$ )	Eq. time (Gyr)	Dyn. time (Gyr)	EQ?
Center $\sim 10^{-2}$	$\sim 10^{-2}$	$\sim 1$	○
Outskirts $\sim 10^{-4}$	$\sim 1$	$\sim 1$	△
Shock $\sim 10^{-2}$	$\sim 10^{-2}$	$\sim 10^{-2*}$	△

\*Sound crossing time through 10 kpc

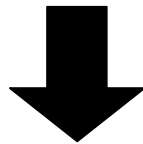
“Equilibrium” is not always a reasonable assumption at the outskirts of clusters!!



### 3) Motivation

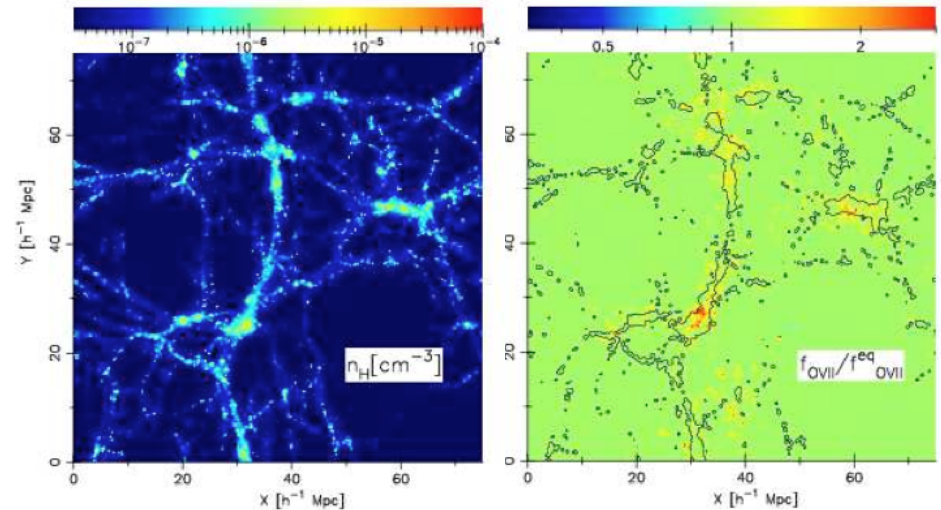


- Actually, in the warm-hot intergalactic medium (WHIM), where the density is lower, a non-EQ state or 2T structure have been pointed out

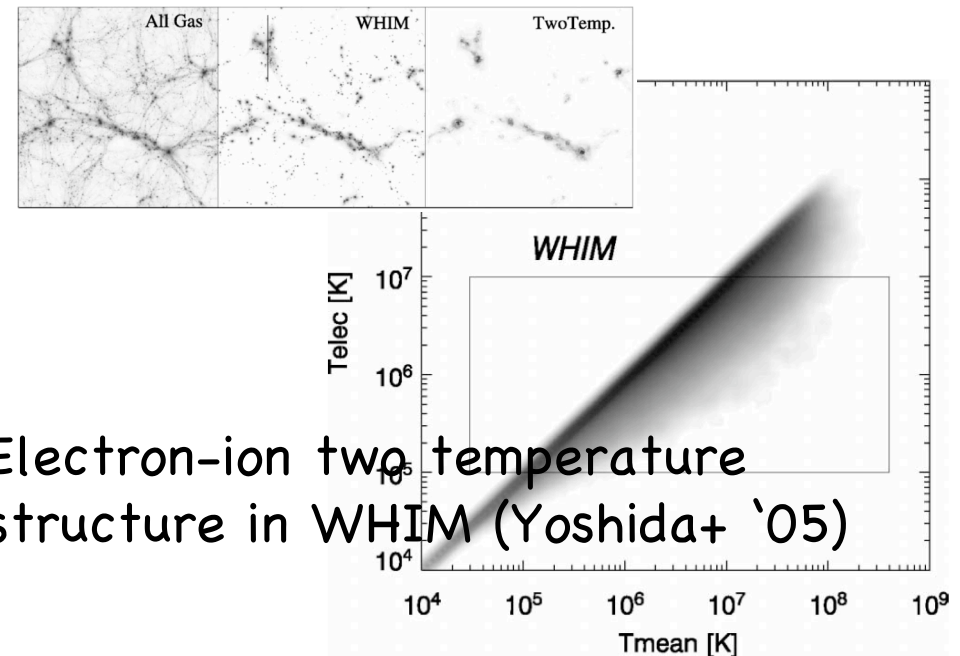


How significant in ICM  
in the linked region of  
the merging cluster?

→ 3D Gas/DM simulations +  
non-EQ & 2T calculations  
*Simultaneously!!*



A non-equilibrium ionization state of O<sub>vii</sub> in WHIM (Yoshikawa, Sasaki '06)



Electron-ion two temperature structure in WHIM (Yoshida+ '05)

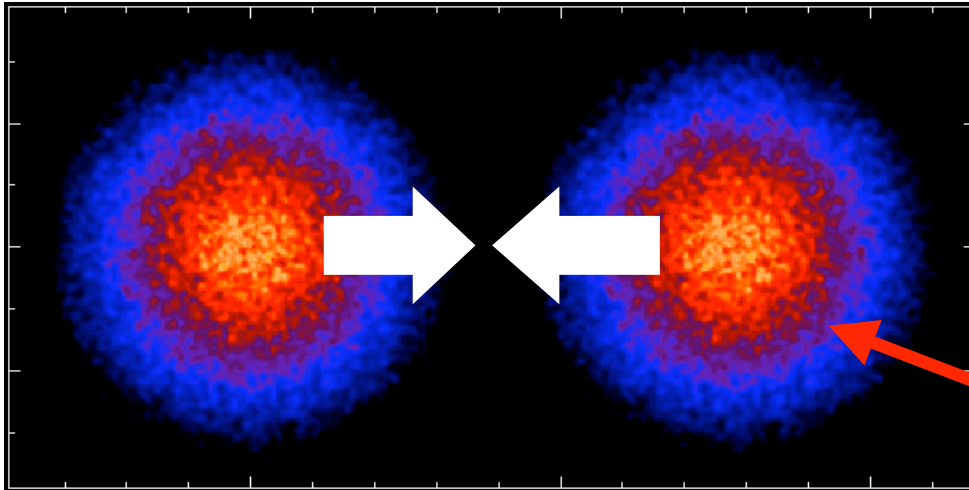


## 4) Model

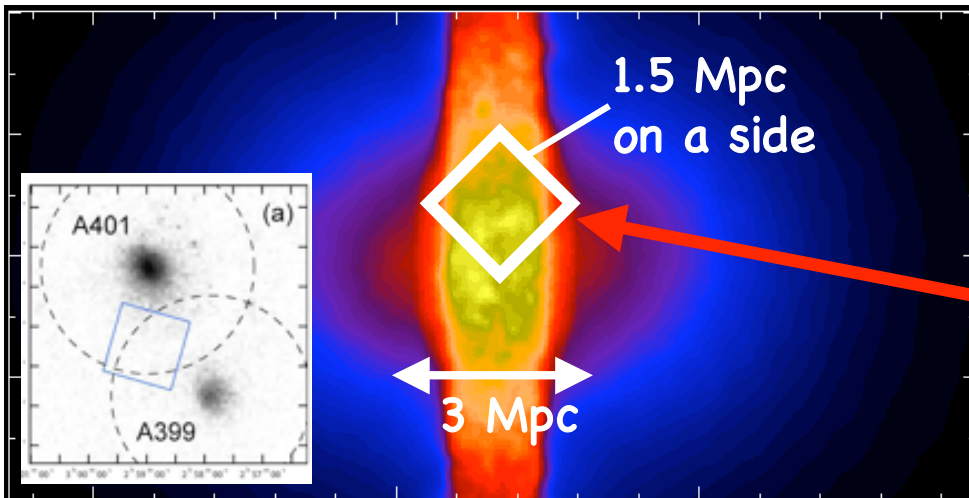


Initial

Initial Property	Value
$n_0^{\text{fit}} (\times 10^{-3} \text{ cm}^{-3})$ , $r_c^{\text{fit}} (\text{Mpc})$ , $\beta^{\text{fit}}$	5.00, 0.16, 0.61
$T_{\text{sl}} (\text{keV})$ at 0, 0.75, 1.5 Mpc	7.64, 5.00, 3.42



Present



To reproduce the situation of A399/401, ...

- 1:1 head-on merger
- face-on view
- Init. ICM distribution
  - Asca: temp. map  
 $T(r) = (T_{\text{vir}}/\beta) \exp(-r/r_{\text{vir}})$
  - Newton:  $\beta$ -model fit  
 $n(r) = n_0 \{1 + (r/r_c)^2\}^{-3\beta/2}$
  - Hydrostatic Equilibrium
- Init. relative velocity
  - 1050 km/s
  - to reproduce the observed temperature of the linked region by  $T_{\text{sl}}^*$

\* $T_{\text{sl}}$ : the spectroscopic-like temperature (Mazzotta et al. 2004)





## 5) Method



- SPH + N-body
  - $10^6$  particles for each component
  - $h \sim 0.1$  Mpc @ the linked region
  - MG+Balsara viscosity
  - FIRST simulator @ CCS, Tsukuba



FIRST simulator  
 496 Host CPUs  
 -3.1Tflops  
 240 Blade-GRAPES  
 -33Tflops  
 1.6TB memory  
 22TB storage

- Ionization fractions
  - Yoshikawa, Sasaki '06
  - SPEX 1.10/CUBA UVX bkg.

$$\frac{df_j}{dt} = \sum_{k=1}^{j-1} S_{j-k,k} f_k - \sum_{i=j+1}^{Z+1} S_{i-j,j} f_j - \alpha_j f_j + \alpha_{j+1} f_{j+1}$$

- E-I two-temperature
  - Takizawa '99

$$\frac{d\tilde{T}_e}{dt} = \frac{\tilde{T}_i - \tilde{T}_e}{t_{ei}} - \frac{\tilde{T}_e}{u} Q_{sh}$$

$$\tilde{T}_i \equiv T_i/T \quad \tilde{T}_e \equiv T_e/T$$

$$T \equiv (n_e T_e + n_i T_i) / (n_e + n_i)$$

$$t_{ei} = 2 \times 10^8 \text{ yr} \frac{(T_e/10^8 \text{ K})^{3/2}}{(n_i/10^{-3} \text{ cm}^{-3})} \cdot \left( \frac{40}{\ln \Lambda} \right)$$

$f_j$ : ionization fraction of  $j$ -times ionized ion,  
 $Z$ : atomic number of H, He, C, N, O, Ne, Mg, Si, S, Fe,  
 $S_{i,j}$ : Ionization rate that ion  $j$  ejects  $i$  electrons,  
 $\alpha_j$ : Recombination rate for ion  $j$

$T_e, n_e$ : the electron temperature and number density,  
 $T_i, n_i$ : the ion temperature and number density,  
 $T, n$ : the gas (mean) temperature and number density,  
 $u$ : the internal energy of the gas,  
 $Q_{sh}$ : the shock (viscosity) heating rate,  
 $t_{ei}$ : the Coulomb scattering timescale



## 6) Result 1: Density, Temperature

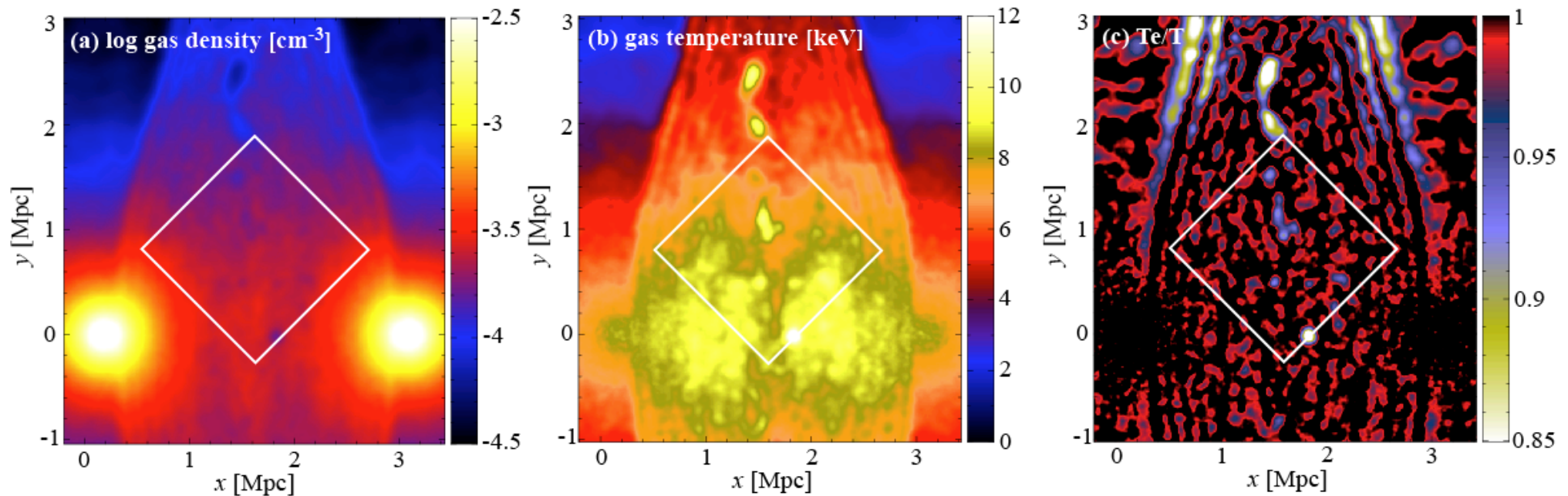


Fig.) The cross sections along the collision axis: the gas density [ $\log \text{cm}^{-3}$ ], the gas mean temperature [keV], the electron temperature/the mean temperature

- The gas density  $\doteq$  observation ( $3 \times 10^{-4} \text{cm}^{-3}$ ) Agreement
- $T$  (8keV)  $>$   $T_{\text{sl}}^*$  (6.5keV)  $\doteq$  observation (6–7keV) Agreement
- No significant shocks in the linked region

↪ Te is only a few percent lower than T

\* $T_{\text{sl}}$ : the spectroscopic-like temperature toward the linked region



## 7) Result 2: Fe Ionization State

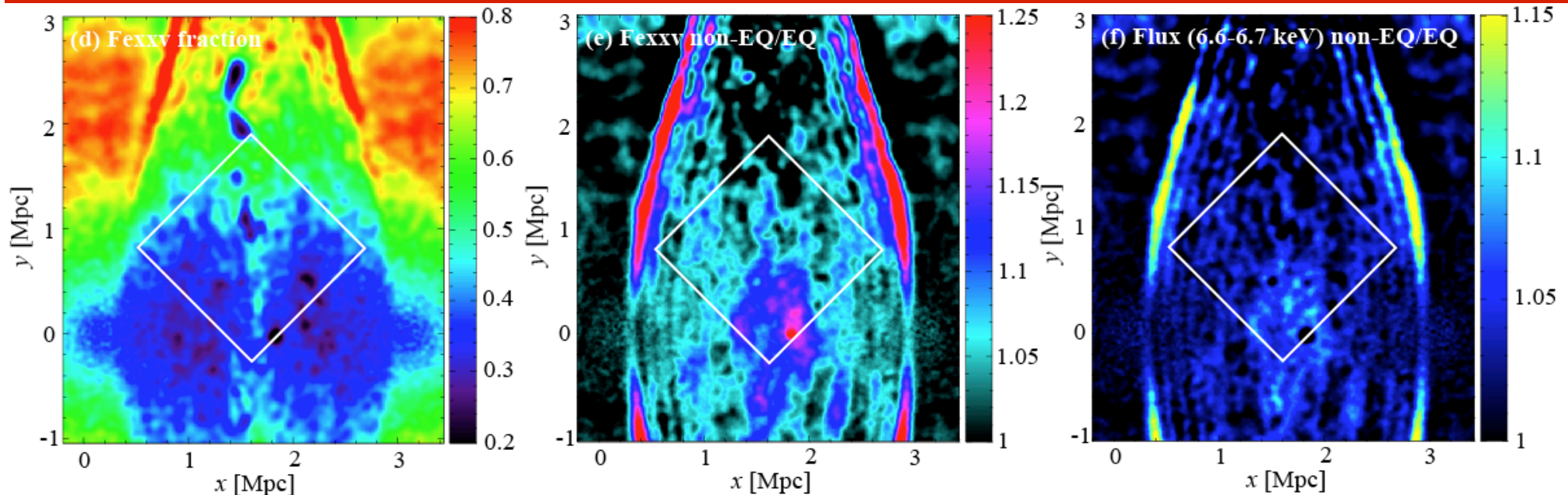


Fig.) The cross sections along the collision axis: the Fexxv fraction, the ratio of the Fexxv fraction relative to that in the equilibrium state, the ratio of the line intensity of 6.6–6.7 keV band relative to that in the equilibrium state.

- Fexxv (He-like) fraction: 30–60%, the largest fraction

Fexxv is 10–20% larger

6.6–6.7keV band

intensity is 5–10% larger



than that in the  
equilibrium state  
(IE & IT)

★See discussion 1!!





## 8) Result 3: Significance in obs.

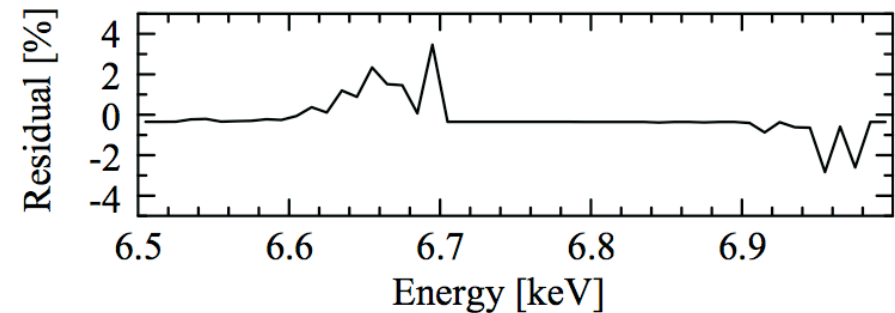
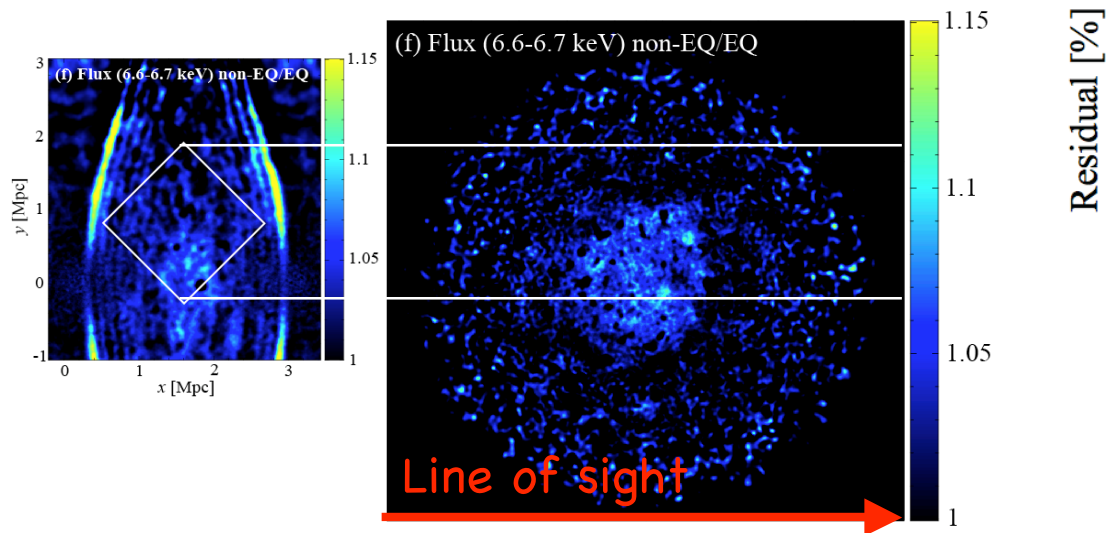


Fig.) The ratio of X-ray spectrum integrated along the line of sight toward the linked region relative to that in the equilibrium state

Fig.) The cross section perpendicular to the collision axis ( $x=1.6\text{Mpc}$ ): the ratio of the line intensity of 6.6–6.7 keV band relative to that in the equilibrium state

- The deviation from the equilibrium state is only significant at around the center of the linked region
- Non-EQ effects on the X-ray spectrum is deluted in integrating along the line of sight

The integrated intensity of 6.6–6.7keV band is only a few percent higher than that in the equilibrium state

★See discussion 2!!



## 9) Discussion 1

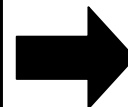


- In the linked region,  $T_e$  increases from 3keV to 8keV according to the adiabatic compression
- $F_{\text{xxv}}$  peaks at 3keV and decreases as  $T_e$  increases

However,  $F_{\text{xxv}}$  fraction is higher than that in the equilibrium state because the ionization of  $F_{\text{xxv}}$  is not fast enough to catch up with the ionization equilibrium state

Thus, the  $F_{\text{xxv}}$  fraction is 10-20% larger than that in the equilibrium state. Meanwhile, the  $F_{\text{xxvi}}$  fraction is 5% smaller.

Line emissions by **collisional excitation and recombination of  $F_{\text{xxv}}$**  are stronger, while that by **recombination of  $F_{\text{xxvi}}$**  is dimmer



As a result, 6.6-6.7keV band intensity is enhanced (by 5-10%)

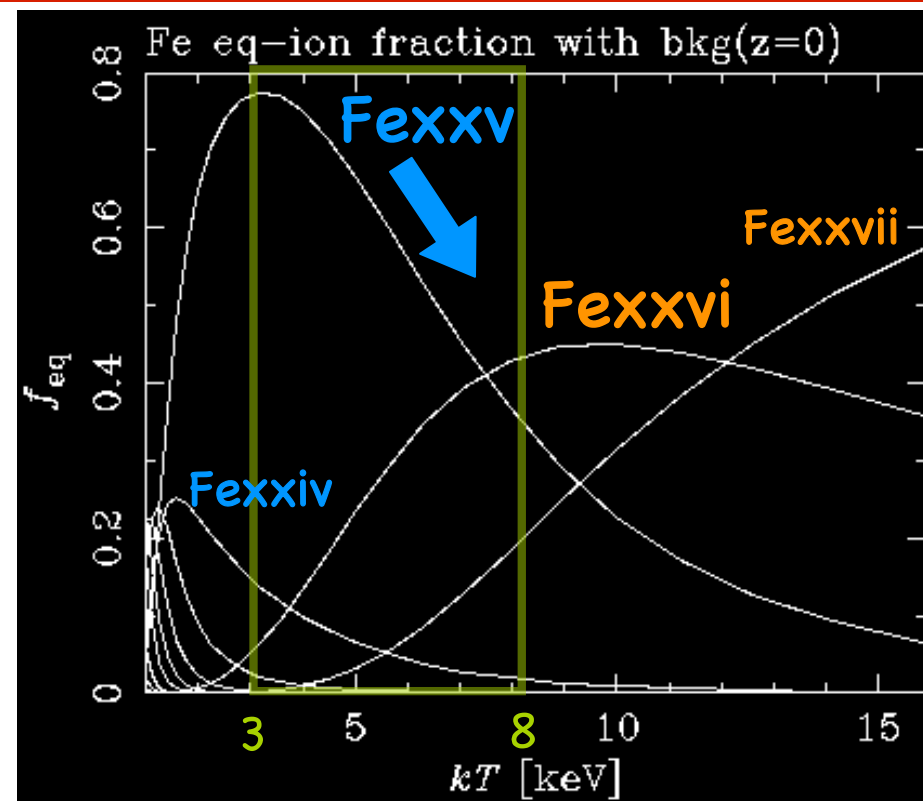


Fig.) ionization fractions of iron in the equilibrium state as a function of  $T_e (=T)$



## 10) Discussion 2



- Metallicity is evaluated from the line intensity, and it would be overestimated in proportion to the intensity enhancement

The enhancement of a few percent is, however, within the uncertainty of the Suzaku observation

Because there is no enhancement of the number density of galaxies (Sakelliou, Ponman '04), and no noticeable small-scale structures (Fujita+ '08), it may be unlikely that the metallicity is concentrated near the center of the linked region.

If the metallicity is concentrated near the center of the linked region, the projection effect is insignificant, though the deviation of the intensity is 10-20% at the maximum.

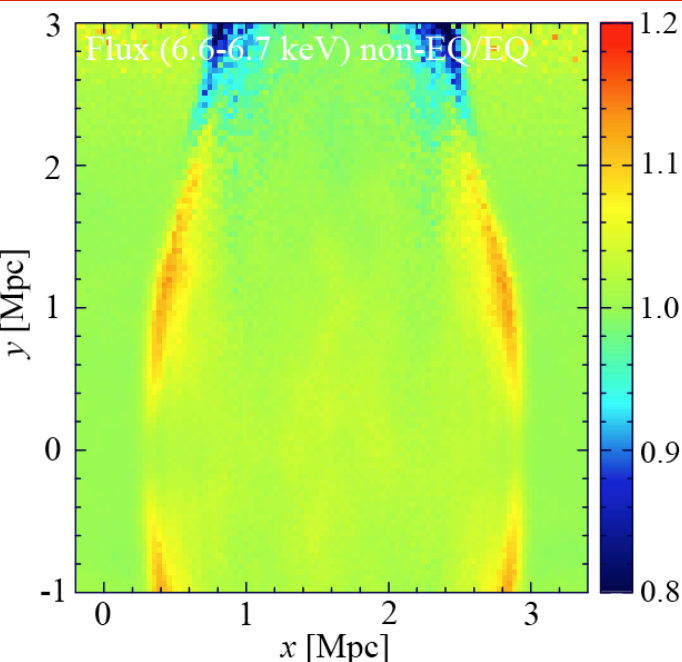
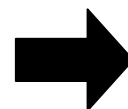


Fig.) The ratio of the intensity of 6.6–6.7 keV band integrated along the line of sight relative to that in the equilibrium state

Therefore, we may say that 0.2 times the solar metallicity estimated by the observation (Fujita et al. 2008) is valid.



# Concluding Remarks



- We investigate a non-equilibrium ionization state and two-temperature structure of the intracluster medium in the linked region of A399/401, using a series of N-body + SPH simulations
- We find in the linked region that:
  1. an ionization fraction of Fe xxv is 10-20 % larger than that in an ionization equilibrium state
  2. the electron temperature is a few percent smaller than the mean temperature of ions and electrons
  3. Fe K emission line intensity integrated along the line of sight is enhanced by a few percent compared with that calculated under the assumptions of the equilibrium state
- HOWEVER, such enhancement is within the uncertainty of the Suzaku observation, suggesting the Suzaku detection of 0.2 times the solar metallicity is not severely affected by a non-EQ ionization state and 2T structure in the region

-future work- We are now doing a systematic study of a non-equilibrium ionization state and two temperature structure in various types (e.g., head-on/off-set, major/minor, ...) of merging clusters. We are also planning a study of the Bullet cluster including radiative cooling.

